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A Comparative Study of Some Spatial-Temporal Discretization Schemes for Nonlinear Magnetohydrodynamic Simulation of Plasmas

D. Nath, M. S. Kalra and P. Munshi

Nuclear Engineering and Technology Program Indian Institute of Technology Kanpur Kanpur - 208 016, India (dnath@iitk.ac.in)



Outline

- Introduction
- MHD Model of plasma
 - Ideal MHD
 - One-dimensional Ideal MHD equations
 - Initial and Boundary Conditions
- Relevant Numerical Schemes and their Stabilities
- Numerical Results and Discussion
- Conclusion



Introduction

Plasmas



- Plasmas are ubiquitous in nature having different densities and temperatures
- Fusion plasmas are required to have high temperature $(>10^8 \text{ K} \sim 10 \text{ keV})$

Fusion Plasma





- Confining plasma using magnetic field in a torus i.e. the Tokamak, is the main candidate for achieving fusion on earth
- Due to collisionless nature, kinetic model (Boltzmann Equation) represent the correct behavior of the plasma
- However fluid model are as applicable in a limited time scales (~ ms) and length scales



Ideal Magnetohydrodynamic Equations

• Continuity equation:

 $\frac{d\rho}{dt} + \rho \nabla \cdot \boldsymbol{v} = 0$

• Momentum equation:

$$\rho \frac{d\boldsymbol{\nu}}{dt} + \nabla \left(P + \frac{B^2}{8\pi} \right) = \frac{\boldsymbol{B} \cdot \nabla \boldsymbol{B}}{8\pi}$$

$$\frac{dI}{dt} + \gamma \nabla \cdot \boldsymbol{v} = 0$$

material derivative: $\frac{d}{dt} = \boldsymbol{v} \cdot \nabla + \frac{\partial}{\partial t}$

• Induction equation:

$$\frac{d\boldsymbol{B}}{dt} - \boldsymbol{B} \cdot \nabla \boldsymbol{v} + \boldsymbol{B} \nabla \cdot \boldsymbol{v} = \begin{cases} \frac{c^2 \eta}{4\pi} \nabla^2 \boldsymbol{B} & : & \text{RMHD} \\ 0 & : & \text{IMHD} \end{cases}$$

dP

One-dimensional MHD Eqns

- MHD forms a hyperbolic system of 8 equations:
 - ► 2 scalar, 3+3 vector equations;
 - ▶ 8 variables: ρ , v_r , v_θ , v_z , P, B_r , B_θ , B_z
- In cylindrical coordinates 1-D reduction:
 - ► *B* is along *z*-direction
 - $B_r = 0, B_\theta = 0 \text{ and } B_z = B$



- assuming no angular rotation in the plasma ($v_{\theta} = 0$)
- no motion along z-direction ($v_z = 0$) so remaining velocity is only along r-direction ($v_r = v$)
- due to symmetry in θ and z directions:

$$\frac{\partial}{\partial \theta} = 0$$
 and $\frac{\partial}{\partial z} = 0$

Problem is reduced into one dimensional in cylindrical coordinates



One-dimensional MHD Equations

- One-dimensional MHD equations in Cylindrical coordinates
- Non-dimensionalzation to reach in Alfvén time scale

$$\begin{aligned} r \to ra, \ t \to ta/v_A, \ \rho \to \rho\rho_o, \ v \to vv_A, \ P \to PP_o, \ B \to BB_a \\ \text{Alfvén Speed} \ v_A &= \sqrt{(B_a^2/\mu_0\rho_o)}, \qquad \beta = P_o/(B_0^2/2\mu_0) \end{aligned}$$
$$\begin{aligned} \frac{\partial \rho}{\partial t} &= -\frac{1}{r}\frac{\partial}{\partial r}(r\rho v) \\ \frac{\partial v}{\partial t} &= -v\frac{\partial v}{\partial r} - \frac{1}{\rho}\frac{\partial}{\partial r}\left(P + \frac{B^2}{2\mu_0}\right) \\ \frac{\partial P}{\partial t} &= -v\frac{\partial P}{\partial r} - \gamma \frac{P}{r}\frac{\partial}{\partial r}(rv) \\ \frac{\partial P}{\partial t} &= -v\frac{\partial P}{\partial r} - \gamma \frac{P}{r}\frac{\partial}{\partial r}(rv) \\ \frac{\partial B}{\partial t} &= -\frac{B}{r}\frac{\partial}{\partial r}(rv) - v\frac{\partial B}{\partial r} \end{aligned}$$

Matrix Form of One-dimensional MHD Equations

• Ideal MHD set of equations can be written in matrix form

$$\frac{\partial \mathbf{U}}{\partial t} + \mathbf{A}(\mathbf{U})\frac{\partial \mathbf{U}}{\partial r} = \mathbf{S}(\mathbf{U}, r) \qquad \mathbf{U} = \begin{cases} \rho \\ v \\ P \\ B \end{cases} \qquad \mathbf{S}(\mathbf{U}, r) = \begin{cases} -\rho v/r \\ 0 \\ -\gamma vP/r \\ -vB/r \end{cases}$$

• Matrix **A**(**U**) and its eigen values:

$$\mathbf{A}(\mathbf{U}) = \begin{bmatrix} v & \rho & 0 & 0 \\ 0 & v & \beta/2\rho & B/\rho \\ 0 & \gamma P & v & 0 \\ 0 & B & 0 & v \end{bmatrix} \qquad \Lambda \equiv |\lambda_A|_{max} = \sqrt{\frac{B^2}{\rho} + \frac{\beta\gamma P}{2\rho}} + |v|$$

• Positive roots implies the hyperbolic nature





Initial and Boundary Conditions

• Initial Conditions

$$B(r,0) = \sqrt{1 - \beta(1 - 3r^2 + 2r^3)}$$

$$P(r,0) = (1 - 3r^2 + 2r^3)[1 + \delta \sin(2\pi r)]$$
 for $0 \le r \le 1$
Constant $\delta = 0.1$

Boundary Conditions

$$r = 0, \qquad \frac{\partial \rho}{\partial r} = 0, \qquad \nu = 0, \qquad \frac{\partial P}{\partial r} = 0, \qquad \frac{\partial B}{\partial r} = 0$$
$$r = 1, \qquad \rho = 0, \qquad \nu = 0, \qquad P = 0, \qquad B = 1$$



Numerical Schemes

- We have compared the stability of three numerical schemes in one dimension in cylindrical coordinates
 - One dimensional MHD equations
 - Cylindrical geometry

 - **Explicit time stepping**
- The Lax-Wendroff-Retchmyer
- MacCormack
- Runge-kutta fourth order

Numerical Schemes



Lax-Wendroff-Retchmyer Scheme

PDE:

$$\frac{\partial \boldsymbol{U}}{\partial t} + \frac{\partial \boldsymbol{F}}{\partial x} = 0$$

Discretization:





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MacCormack Scheme

PDE:

 $\frac{\partial \boldsymbol{U}}{\partial t} + \frac{\partial \boldsymbol{F}}{\partial x} = 0$





Runge-Kutta Fourth Order





- Semi-Discritization for spatiotemporal PDE (Method of Lines):
 - First discretized in space, obtain ODEs in time

 $\frac{dU_j(t)}{dt} = f(t, F_j)$

□ Apply RK4 to advance in time

$$k^{n1} = f(t^{n}, F_{j}^{n})$$

$$k^{n2} = f\left(t^{n} + \frac{1}{2}\Delta t, F_{j}^{n} + \frac{1}{2}\Delta t k^{n1}\right)$$

$$k^{n3} = f\left(t^{n} + \frac{1}{2}\Delta t, F_{j}^{n} + \frac{1}{2}\Delta t k^{n2}\right)$$

$$k^{n4} = \left(t^{n} + \Delta t, F_{j}^{n} + \Delta t k^{n3}\right)$$

$$U^{n+1} = U^{n} + \frac{\Delta t}{6}(k^{n1} + 2k^{n2} + 2k^{n3} + k^{n4})$$

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Stability Criteria

• A scalar linear advection equation with advection speed ∧:

$$\frac{\partial u}{\partial t} + \Lambda \frac{\partial u}{\partial r} = 0$$

- Von-Neumann stability analysis: $u_k(x,t) = \hat{u}_k e^{i\phi}$ where $\phi = \omega \Delta t = k \frac{\Delta x}{\Delta t} \Delta t$
- **Numerical amplification Factor:** $|G| = \frac{u_k^{n+1}}{u_k^n} \le 1$ for stability
- Lax-Wendroff-Retchmyer or MacCormak

$$G = \left(1 - \frac{1}{2}C^{2}\sin^{2}\phi + \frac{1}{24}C^{4}\sin^{4}\phi\right) + j\left(-C\sin\phi + \frac{1}{6}C^{3}\sin^{3}\phi\right)$$
$$|G| = \sqrt{1 - 4C^{2}(1 - C^{2})\sin^{4}(\phi/2)}$$
$$C = \frac{1}{2}C^{2}$$

 $C = \frac{\Lambda \Delta t}{\Delta r} \le 1$

• Runge-Kutta fourth order

$$G = (1 - C^{2} + C^{2} \cos \phi) + j(-C \sin \phi); \quad j = \sqrt{-2}$$
$$|G| = \sqrt{1 - \frac{1}{72}}C^{6} \left(1 - \frac{1}{8}C^{2} \sin^{2} \phi\right) \sin^{6} \phi$$

$$C = \frac{\Lambda \cdot \Delta t}{\Delta r} \le 2\sqrt{2}$$

Numerical Results

- Effect of Schemes
- Effect of Courant Number
- Effect of Grid Refinement
- Long Time Run
- Stability





Effect of Schemes



- LWR, MAC and RK4 schemes can advance more than 4 Alfven time scales for grid points 61 and r = 0.25
- LWR suffers from oscillations comparatively earlier and reaches high around time ~ $4 \tau_a$
- Similar behavior is observed in time stepping of Magnetic field



Effect of Courant Number



- Keeping fixed grid and changing Courant Number from C = 0.2 to C = 0.8 LWR suffers from oscillation from the early phases
- MAC or RK4 are smoother





Effect of Grid Refinement



- On increasing the grid points again LWR surprisingly have lot of numerical instabilities whereas MAC and RK4 have reached grid independence.
- C = 0.8 (Kept Fixed)





Long Time Run

• From t = 5 to 15, burst of instabilities appear in LWR (C = 0.2, Grid points = 61)



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Radial Distribution



• A quasi-steady state is reached after t=50

Stability



Stability and Courant Number



- The growth of numerical instabilities begins from the edge r = 1. At r = 1 Courant number become high, and if comes out from the stability region, makes an scheme numerically unstable
- Stability regions of LWR (with C = 1.0, 0.8) and RK4 ($C = 2\sqrt{2}$)
- Any scheme to be stable, needs to be operated within its stability region



Numerical Amplification and Phase Error in LWR



- Numerical amplification variation w.r.t the mesh wave number is negligible for Courant numbers near to C = 1.0 and C = 0.1 and maximum for C = 0.8
- Phase error for C = 0.8 become negative shows numerical instability in the scheme
- A positive dispersion (phase) error implies that numerical advection velocity is larger than the exact physical velocity and vice versa



Numerical Amplification and Phase Error in RK4



- Less numerical dissipation in RK4 comparatively LWR and MAC
- Phase error is almost independent of the Courant number
- Phase error continuously increases towards the high wave numbers



Conclusion

- Three widely used discretization schemes applied to a representative *nonlinear* problem arising in the MHD simulation of plasmas: LWR, MAC and RK4
- LWR and RK4 are not suitable schemes for nonlinear hyperbolic MHD equation
- MAC scheme always gave stable results if CFL is satisfied
- The linear stability conditions are **necessary** but **not sufficient** to guarantee the numerical stability of these algorithms when applied to a *nonlinear* **hyperbolic problem**
- The solution for the transient phase obtained from the three schemes are found to differ significantly due to different amount of **numerical diffusion and dispersion** present in each scheme



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Thank you!

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